# SINGULARITIES OF GENERIC LIGHTCONE GAUSS MAPS AND LIGHTCONE PEDAL SURFACES OF SPACELIKE CURVES IN MINKOWSKI 4-SPACE 

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#### Abstract

The main goal of this paper is to study singularities of generic lightcone Gauss maps and lightcone pedal surfaces of spacelike curves in Minkowski 4-space. To do this, we first construct lightcone height functions and extended lightcone height functions, and then show the relations between singularities of generic lightcone Gauss maps(resp. lightcone pedal surfaces) and that of lightcone height functions(resp. extended lightcone height functions). In addition some geometric properties of the spacelike curves are studied from geometrical point of view.


## 1. Introduction

The study of Minkowski 4 -space has produced fruitful results, see for example $[5,6,7,8]$. Motivated by the study of generic lightcone Gauss maps of spacelike curves in Minkowski 4 -space, and completing the study of submanifolds in Minkowski 4 -space from the singularity theory point of view, we develop the study of singularities of generic lightcone Gauss maps and lightcone pedal surfaces of spacelike curves in Minkowski 4-space. To do this we need to work out local differential geometry tools for the spacelike curves similar to those of curves in Euclidean space[1, 2, 3, 4]. As it was expected, the situation presents certain peculiarities when it is compared with the Euclidean space and the Minkowski 3 -space. For instance, the dimension of lightlike normal vector space is 2 , then

[^0]there exist many lightcone Gauss maps and lightcone pedal curves. For some basic notions in Lorentzian geometry, see [9].

The paper is organized as follows:
The rest of this section introduces the basic notions of Minkowski 4-space, after these basic notions, we give the local differential geometry of spacelike curves and the main results of this paper. Section 2 first constructs the lightcone height functions, which are useful tools for the study of singularities of lightcone Gauss maps, and then shows the relations between singularities of lightcone Gauss maps and that of lightcone height functions. Sections 3 and 4 give the proofs of the main results of this paper.

Let $\mathbb{R}^{4}=\left\{\left(x_{1}, x_{2}, x_{3}, x_{4}\right) \mid x_{1}, x_{2}, x_{3}, x_{4} \in \mathbb{R}\right\}$ be a 4 -dimensional vector space. For any two vectors $\boldsymbol{x}=\left(x_{1}, x_{2}, x_{3}, x_{4}\right), \boldsymbol{y}=\left(y_{1}, y_{2}, y_{3}, y_{4}\right)$ in $\mathbb{R}^{4}$, the pseudo scalar product of $\boldsymbol{x}$ and $\boldsymbol{y}$ is defined by $\langle\boldsymbol{x}, \boldsymbol{y}\rangle=-x_{1} y_{1}+x_{2} y_{2}+x_{3} y_{3}+x_{4} y_{4}$. We call $\left(\mathbb{R}^{4},\langle\rangle,\right)$ a Minkowski 4 -space and denote it by $\mathbb{R}_{1}^{4}$.

We say that a vector $\boldsymbol{x}$ in $\mathbb{R}_{1}^{4} \backslash\{0\}$ is a spacelike vector, a lightlike vector or a timelike vector if $\langle\boldsymbol{x}, \boldsymbol{x}\rangle$ is positive, zero, negative respectively. For a vector $\boldsymbol{n} \in \mathbb{R}_{1}^{4}$ and a real number $c$, a hyperplane with pseudo normal $\boldsymbol{n}$ is defined by $L H P(\boldsymbol{n}, c)=\left\{\boldsymbol{x} \in \mathbb{R}_{1}^{4} \mid\langle\boldsymbol{x}, \boldsymbol{n}\rangle=c\right\}$. The norm of a vector $\boldsymbol{x} \in \mathbb{R}_{1}^{4}$ is defined by $\|\boldsymbol{x}\|=\sqrt{|\langle\boldsymbol{x}, \boldsymbol{x}\rangle|}$. For any two vectors $\boldsymbol{x}, \boldsymbol{y}$ in $\mathbb{R}_{1}^{4}$, we say that $\boldsymbol{x}$ is pseudoperpendicular to $\boldsymbol{y}$ if $\langle\boldsymbol{x}, \boldsymbol{y}\rangle=0$.

An orientation and a timelike orientation of $\mathbb{R}_{1}^{4}$ are fixed(i.e., a 4 -volume form $d V$, and future time-like vector field, have been chosen). Let $\gamma: I \rightarrow \mathbb{R}_{1}^{4}$ be a smooth regular curve in $\mathbb{R}_{1}^{4}$ (i.e., $\dot{\gamma}(t) \neq 0$ for any $t \in I$ ), where $I$ is an open interval. We say that a smooth regular curve $\gamma$ is a spacelike curve if $\langle\dot{\gamma}(t), \dot{\gamma}(t)\rangle>0$ for any $t \in I$. The arclength of a spacelike curve $\gamma$, measured from $\gamma\left(t_{0}\right)\left(t_{0} \in I\right)$, is

$$
s(t)=\int_{t_{0}}^{t}\|\dot{\gamma}(t)\| d t
$$

Then a parameter $s$ is determined such that $\left\|\gamma^{\prime}(s)\right\|=1$, where $\gamma^{\prime}(s)=d \gamma / d s(s)$. Consequently we say that a spacelike curve $\gamma$ is parameterized by arclength if $\left\|\gamma^{\prime}(s)\right\|=1$. Throughout the rest of this paper $s$ is assumed arclength parameter. Let $\boldsymbol{t}(s)$ denote $\gamma^{\prime}(s)$. We call $\boldsymbol{t}(s)$ a unit tangent vector of $\gamma$ at $s$. The signature of $\boldsymbol{x}$ is defined to be

$$
\delta(\boldsymbol{x})=\operatorname{sign}(\boldsymbol{x})= \begin{cases}1 & \boldsymbol{x}: \text { spacelike } \\ 0 & \boldsymbol{x}: \text { lightlike } \\ -1 & \boldsymbol{x}: \text { timelike }\end{cases}
$$

For any $\boldsymbol{x}_{1}, \boldsymbol{x}_{2}, \boldsymbol{x}_{3} \in \mathbb{R}_{1}^{4}$, we define a vector $\boldsymbol{x}_{1} \wedge \boldsymbol{x}_{2} \wedge \boldsymbol{x}_{3}$ by

$$
\begin{aligned}
& \boldsymbol{x}_{1} \wedge \boldsymbol{x}_{2} \wedge \boldsymbol{x}_{3}=\left|\begin{array}{cccc}
-e_{1}, & e_{2}, & e_{3}, & e_{4} \\
x_{1}^{1}, & x_{1}^{2}, & x_{1}^{3}, & x_{1}^{4} \\
x_{2}^{1}, & x_{2}^{2}, & x_{2}^{3}, & x_{2}^{4} \\
x_{3}^{1}, & x_{3}^{2}, & x_{3}^{3}, & x_{3}^{4}
\end{array}\right| \\
& =\left(-\left|\begin{array}{lll}
x_{1}^{2}, & x_{1}^{3}, & x_{1}^{4} \\
x_{2}^{2}, & x_{2}^{3}, & x_{2}^{4} \\
x_{3}^{2}, & x_{3}^{3}, & x_{3}^{4}
\end{array}\right|,-\left|\begin{array}{lll}
x_{1}^{1}, & x_{1}^{3}, & x_{1}^{4} \\
x_{2}^{1}, & x_{2}^{3}, & x_{2}^{4} \\
x_{3}^{1}, & x_{3}^{3}, & x_{3}^{4}
\end{array}\right|,\left|\begin{array}{lll}
x_{1}^{1}, & x_{1}^{2}, & x_{1}^{4} \\
x_{2}^{1}, & x_{2}^{2}, & x_{2}^{4} \\
x_{3}^{1}, & x_{3}^{2}, & x_{3}^{4}
\end{array}\right|,-\left|\begin{array}{lll}
x_{1}^{1}, & x_{1}^{2}, & x_{1}^{3} \\
x_{2}^{1}, & x_{2}^{2}, & x_{2}^{3} \\
x_{3}^{1}, & x_{3}^{2}, & x_{3}^{3}
\end{array}\right|\right)
\end{aligned}
$$

where $\boldsymbol{x}_{i}=\left(x_{i}^{1}, x_{i}^{2}, x_{i}^{3}, x_{i}^{4}\right)$.
For a unit speed spacelike curve $\gamma: I \rightarrow \mathbb{R}_{1}^{4}$ with $\left\|\gamma^{\prime \prime}(s)\right\| \neq 0$ and $\left\|\boldsymbol{n}_{1}^{\prime}(s)+\delta_{1} k_{1}(s) \boldsymbol{t}(s)\right\| \neq 0$, then we can construct a pseudo-orthogonal frame $\left\{\boldsymbol{t}(s), \boldsymbol{n}_{1}(s), \boldsymbol{n}_{2}(s), \boldsymbol{n}_{3}(s)\right\}$, which satisfies the following Frenet-Serret type formulae, of $\mathbb{R}_{1}^{4}$ along $\gamma$.

$$
\begin{cases}\boldsymbol{t}(s) & =\gamma^{\prime}(s) \\ \boldsymbol{t}^{\prime}(s) & =k_{1}(s) \boldsymbol{n}_{1}(s) \\ \boldsymbol{n}_{1}^{\prime}(s) & =-\delta_{1} k_{1}(s) \boldsymbol{t}(s)+k_{2}(s) \boldsymbol{n}_{2}(s) \\ \boldsymbol{n}_{2}^{\prime}(s) & =\delta_{3} k_{2}(s) \boldsymbol{n}_{1}(s)+k_{3}(s) \boldsymbol{n}_{3}(s) \\ \boldsymbol{n}_{3}^{\prime}(s) & =\delta_{1} k_{3}(s) \boldsymbol{n}_{2}(s)\end{cases}
$$

where $k_{1}(s)=\left\|\gamma^{\prime \prime}(s)\right\|, \quad \boldsymbol{n}_{1}(s)=\frac{\gamma^{\prime \prime}(s)}{k_{1}(s)}, \quad \boldsymbol{n}_{2}(s)=\frac{\boldsymbol{n}_{1}^{\prime}(s)+\delta_{1} k_{1}(s) \boldsymbol{t}(s)}{\left\|\boldsymbol{n}_{1}^{\prime}(s)+\delta_{1} k_{1}(s) \boldsymbol{t}(s)\right\|}$, $\boldsymbol{n}_{3}(s)=\boldsymbol{t}(s) \wedge \boldsymbol{n}_{1}(s) \wedge \boldsymbol{n}_{2}(s)$, and $\delta_{i}=\delta\left(\boldsymbol{n}_{i}(s)\right)(i=1,2,3)$. This is the fundamental formula for the study of generic curves in $\mathbb{R}_{1}^{3}$; It is, however, useless at the point $\gamma(s)$ with $\left\|\boldsymbol{n}_{1}^{\prime}(s)+\delta_{1} k_{1}(s) \boldsymbol{t}(s)\right\|=0$. We now denote $A(s)=\boldsymbol{n}_{1}^{\prime}(s)+\delta_{1} k_{1}(s) \boldsymbol{t}(s)(s)$ and $C(s)=\boldsymbol{t}(s) \wedge \boldsymbol{n}_{1}(s) \wedge A(s)$. If $k_{2}(s)=0$, then $A(s)$ is a lightlike vector, so that any pseudo perpendicular vector in the plane normal to $\boldsymbol{t}$ and $\boldsymbol{n}_{1}$ is parallel to $A(s)$.

On the other hand, there exists a lightlike vector $B(s)$ such that $\langle A(s), B(s)\rangle=$ $1,\langle\boldsymbol{t}(s), B(s)\rangle=0,\left\langle\boldsymbol{n}_{1}(s), B(s)\right\rangle=0$ and $\left\{\boldsymbol{t}(s), \boldsymbol{n}_{1}(s), A(s), B(s)\right\}$ is a basis of $\mathbb{R}_{1}^{4}$.

Let

$$
L C_{p}=\left\{\boldsymbol{x}=\left(x_{1}, x_{2}, x_{3}, x_{4}\right) \in \mathbb{R}_{1}^{4} \mid-\left(x_{1}-p_{1}\right)^{2}+\sum_{i=2}^{4}\left(x_{i}-p_{i}\right)^{2}=0\right\}
$$

and

$$
S_{+}^{2}=\left\{\boldsymbol{x} \in L C_{0} \mid \boldsymbol{x}=\left(1, x_{2}, x_{3}, x_{4}\right)\right\}
$$

where $p=\left(p_{1}, p_{2}, p_{3}, p_{4}\right) \in \mathbb{R}_{1}^{4}$. We call $L C_{p}^{*}=L C_{p} \backslash\{p\}$ a lightcone at the vertex $p$ and $S_{+}^{2}$ a lightlike unite sphere. If $\boldsymbol{x}=\left(x_{1}, x_{2}, x_{3}, x_{4}\right)$ is in $L C_{0}^{*}$, we then have $\widetilde{\boldsymbol{x}}=\left(1, x_{2} / x_{1}, x_{3} / x_{1}, x_{4} / x_{1}\right) \in S_{+}^{2}$.

If $k_{2}(s) \neq 0$, for any vector $\boldsymbol{v}=\left(\lambda_{1} \boldsymbol{n}_{1}+\mu_{1} \boldsymbol{n}_{2}+\eta_{1} \boldsymbol{n}_{3}\right)(s) \in S_{+}^{2} \cap N_{s} \gamma$, then

$$
\left\{\begin{array}{l}
\delta_{1} \lambda_{1}^{2}+\delta_{2} \mu_{1}^{2}+\delta_{3} \eta_{1}^{2}=0 \\
\lambda_{1} \boldsymbol{n}_{11}(s)+\mu_{1} \boldsymbol{n}_{21}(s)+\eta_{1} \boldsymbol{n}_{31}(s)=1
\end{array}\right.
$$

Thus locally there exist $f_{1}$ and $g_{1}$ such that $\boldsymbol{v}=\left( \pm \lambda_{1} \boldsymbol{n}_{1}+f\left(\lambda_{1}\right) \boldsymbol{n}_{2}+g\left(\lambda_{1}\right) \boldsymbol{n}_{3}\right)(s)$. where $\lambda_{1} \in \mathbb{R}, \boldsymbol{n}_{a}=\left(n_{a 1}, n_{a 2}, n_{a 3}, n_{a 4}\right)(a=1,2,3)$.

If $k_{2}(s)=0$, for any vector $\boldsymbol{v}=\left(\lambda_{2} \boldsymbol{n}_{1}+\mu_{2} A+\eta_{2} B\right)(s) \in S_{+}^{2} \cap N_{s} \gamma$, we have

$$
\left\{\begin{array}{l}
\lambda_{2}^{2}+2 \mu_{2} \eta_{2}=0 \\
\lambda_{2} \boldsymbol{n}_{11}(s)+\mu_{2} A_{1}(s)+\eta_{2} B_{1}(s)=1
\end{array}\right.
$$

Thus locally there exist $f_{2}$ and $g_{2}$ such that $\boldsymbol{v}=\left( \pm \lambda_{2} \boldsymbol{n}_{1}+f_{2}\left(\lambda_{2}\right) A+g_{2}\left(\lambda_{2}\right) B\right)(s)$, where $\lambda_{2} \in \mathbb{R}, \boldsymbol{n}_{1}=\left(n_{11}, n_{12}, n_{13}, n_{14}\right), A=\left(A_{1}, A_{2}, A_{3}, A_{4}\right) B=\left(B_{1}, B_{2}, B_{3}, B_{4}\right)$. When $\boldsymbol{n}_{1}$ is a spacelike vector, for any vector

$$
\boldsymbol{v}_{0}=\left\{\begin{array}{l}
\left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{2}+g_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{3}\right)(s), \quad k_{2}(s) \neq 0 \\
\left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{2}\left(\lambda_{0}\right) A+g_{2}\left(\lambda_{0}\right) B\right)(s), \quad k_{2}(s)=0
\end{array}\right.
$$

in $S_{+}^{2} \cap N_{s} \gamma$, there exists a rotation map $\rho: S_{+}^{2} \cap N_{s} \gamma \rightarrow S_{+}^{2} \cap N_{s} \gamma$ such that $\rho\left(v_{0}\right)=\widetilde{A \pm C}(s)$.

We define surface

$$
L G_{\gamma}^{\sigma}: I \times \mathbb{R} \rightarrow S_{+}^{2}
$$

by

$$
L G_{\gamma}^{\sigma}(s, \lambda)= \begin{cases}\left(\sigma \lambda \boldsymbol{n}_{1}+f_{1}(\lambda) \boldsymbol{n}_{2}+g_{1}(\lambda) \boldsymbol{n}_{3}\right)(s), \quad k_{2}(s) \neq 0 \\ \left(\sigma \lambda \boldsymbol{n}_{1}+f_{2}(\lambda) A+g_{2}(\lambda) B\right)(s), \quad k_{2}(s)=0\end{cases}
$$

and surface

$$
L P_{\gamma}^{\sigma}: I \times \mathbb{R} \rightarrow L C_{0}^{*}
$$

by

$$
L P_{\gamma}^{\sigma}(s, \lambda)=\langle\gamma(s), \rho(\boldsymbol{v}(s, \lambda))\rangle \boldsymbol{v}(s, \lambda)
$$

where

$$
\boldsymbol{v}(s, \lambda)=\left\{\begin{array}{l}
\left(\sigma \lambda \boldsymbol{n}_{1}+f_{1}(\lambda) \boldsymbol{n}_{2}+g_{1}(\lambda) \boldsymbol{n}_{3}\right)(s), \quad k_{2}(s) \neq 0 \\
\left(\sigma \lambda \boldsymbol{n}_{1}+f_{2}(\lambda) A+g_{2}(\lambda) B\right)(s), \quad k_{2}(s)=0
\end{array}\right.
$$

We call $L G_{\gamma}^{\sigma}$ the lightcone Gauss surface of $\gamma, L P_{\gamma}^{\sigma}$ the lightcone pedal surface of $\gamma$. In fact, for any fixed $\eta_{0} \in \mathbb{R}, L G_{\gamma}^{\sigma}\left(-, \eta_{0}\right)$, denoted by $L G_{\gamma, \eta_{0}}^{\sigma}$ is a lightcone Gauss map and $L P_{\gamma}^{\sigma}\left(-, \eta_{0}\right)$, denoted by $L P_{\gamma, \eta_{0}}^{\sigma}$ is a lightcone pedal curve of $\gamma$.

The study of lightcone Gauss surface of $\gamma$ is of course a very interesting aspect of the situation, from which one may deduce deep results concerning the curve
$\gamma$. In the next step, we shall study that surface. Classically, starting with curve theory, we shall consider the lightcone Gauss map in this paper.

Let $F: \mathbb{R}_{1}^{4} \rightarrow \mathbb{R}$ be a submersion and $\gamma$ a unit speed spacelike curve. We say that $\gamma$ and $F^{-1}(0)$ have $k$-point contact at $s_{0}$ provided the function $g(s)=F \circ \gamma(s)$ satisfies $g\left(s_{0}\right)=g^{\prime}\left(s_{0}\right)=\cdots=g^{(k-1)}\left(s_{0}\right)=0$ and $g^{(k)}\left(s_{0}\right) \neq 0$. Dropping the condition $g^{(k)}\left(s_{0}\right) \neq 0$ we say that there is at least $k$-point contact.

Let $\gamma: S^{1} \rightarrow \mathbb{R}_{1}^{4}$ be a spacelike curve with $k_{1}(s) \neq 0$. We consider the following properties of $\gamma$.
$\left(\mathrm{A}_{1}\right)$ The number of points at which $\gamma$ and hyperplane have 4-point contact is finite.
$\left(\mathrm{A}_{2}\right)$ There is no point at which $\gamma$ and hyperplane have at least 5 -point contact. Our main results are the following.
Theorem A Let $\operatorname{Im}\left(S^{1}, \mathbb{R}_{1}^{4}\right)$ be a space of spacelike curves equipped with Whitney $C^{\infty}$-topology. Then the set of spacelike curves that satisfy $\left(A_{1}\right)$ and $\left(A_{2}\right)$ is a residual set in $\operatorname{Im}\left(S^{1}, \mathbb{R}_{1}^{4}\right)$.

cusp

cuspidal edge

swallowtail

Theorem B Under the assumptions of $\left(A_{1}\right)$ and $\left(A_{2}\right)$,
(a) The lightcone Gauss map $L G_{\gamma, \eta_{0}}^{\sigma}$ of $\gamma$ has a cusp point at $s_{0}$ if and only if $k_{2}\left(s_{0}\right)=0$.
(b) The lightcone pedal surface $L P_{\gamma}^{\sigma}$ is locally diffeomorphic to the cuspidal edage if and only if $\gamma$ and hyperplane $\operatorname{LHP}\left(\rho\left(\boldsymbol{v}_{0}^{\sigma}\right), c_{0}^{\sigma}\right)$ have 3-point contact.
(c) The lightcone pedal surface $L P_{\gamma}^{\sigma}$ is locally diffeomorphic to the swallowtail if and only if $k_{2}\left(s_{0}\right)=0$ (or $\gamma$ and hyperplane $\operatorname{LHP}\left(\rho\left(\boldsymbol{v}_{0}^{\sigma}\right), c_{0}^{\sigma}\right)$ have 4-point contact).

Here

$$
\boldsymbol{v}_{0}^{\sigma}=\left\{\begin{array}{l}
\left(\sigma \lambda \boldsymbol{n}_{1}+f_{1}(\lambda) \boldsymbol{n}_{2}+g_{1}(\lambda) \boldsymbol{n}_{3}\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right) \neq 0 \\
\left(\sigma \lambda \boldsymbol{n}_{1}+f_{2}(\lambda) A+g_{2}(\lambda) B\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right)=0
\end{array}\right.
$$

$c_{0}^{\sigma}=\left\langle\gamma\left(s_{0}\right), \rho\left(\boldsymbol{v}_{0}^{\sigma}\right)\right\rangle$ and $\sigma=+$ or - .

## 2. Lightcone height function on spacelike curve in $\mathbb{R}_{1}^{4}$

We first define a function

$$
H: I \times S_{+}^{2} \cap N \gamma \rightarrow \mathbb{R}
$$

by $H(s, \boldsymbol{v})=\langle\gamma(s), \rho(\boldsymbol{v})\rangle$. We call it the lightcone height function of $\gamma$. For any fixed $\boldsymbol{v}_{0} \in S_{+}^{2} \cap N \gamma$, we let $h_{v_{0}}(s)$ denote $H\left(s, \boldsymbol{v}_{0}\right)$. Then we have the following proposition.

Proposition 2.1. Let $\gamma: I \rightarrow \mathbb{R}_{1}^{4}$ be a unit speed spacelike curve with $k_{1}(s) \neq 0$.
Then:
(1) $h_{v_{0}}^{\prime}\left(s_{0}\right)=0$ if and only if $\rho\left(\boldsymbol{v}_{0}\right) \in N_{s_{0}} \gamma$.
(2) $h_{v_{0}}^{\prime}\left(s_{0}\right)=h_{v_{0}}^{\prime \prime}\left(s_{0}\right)=0$ if and only if

$$
\boldsymbol{v}_{0}= \begin{cases}\left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{2}+g_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{3}\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right) \neq 0 \\ \left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{2}\left(\lambda_{0}\right) A+g_{2}\left(\lambda_{0}\right) B\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right)=0\end{cases}
$$

or $\rho\left(v_{0}\right)=\widetilde{B}\left(s_{0}\right)$.
(3) $h_{v_{0}}^{\prime}\left(s_{0}\right)=h_{v_{0}}^{\prime \prime}\left(s_{0}\right)=h_{v_{0}}^{\prime \prime \prime}\left(s_{0}\right)=0$ if and only if

$$
\boldsymbol{v}_{0}=\left\{\begin{array}{l}
\left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{2}+g_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{3}\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right) \neq 0 \\
\left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{2}\left(\lambda_{0}\right) A+g_{2}\left(\lambda_{0}\right) B\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right)=0
\end{array}\right.
$$

and $k_{2}\left(s_{0}\right)=0$.
(4) $h_{v_{0}}^{\prime}\left(s_{0}\right)=h_{v_{0}}^{\prime \prime}\left(s_{0}\right)=\cdots=h_{v_{0}}^{(4)}\left(s_{0}\right)=0$ if and only if

$$
\boldsymbol{v}_{0}= \begin{cases}\left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{2}+g_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{3}\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right) \neq 0 \\ \left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{2}\left(\lambda_{0}\right) A+g_{2}\left(\lambda_{0}\right) B\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right)=0\end{cases}
$$

and $k_{2}\left(s_{0}\right)=\left\langle A^{\prime}\left(s_{0}\right), A\left(s_{0}\right)\right\rangle=0$.
Proof. (1) $h_{v_{0}}^{\prime}\left(s_{0}\right)=0$ if and only if $\left\langle\gamma^{\prime}\left(s_{0}\right), \rho\left(\boldsymbol{v}_{0}\right)\right\rangle=\left\langle\boldsymbol{t}\left(s_{0}\right), \rho\left(\boldsymbol{v}_{0}\right)\right\rangle=0$, which is equivalent to $\rho\left(\boldsymbol{v}_{0}\right) \in N_{s_{0}} \gamma$.
(2) $h_{v_{0}}^{\prime}\left(s_{0}\right)=h_{v_{0}}^{\prime \prime}\left(s_{0}\right)=0$ if and only if $\left\langle\gamma^{\prime}\left(s_{0}\right), \rho\left(\boldsymbol{v}_{0}\right)\right\rangle=\left\langle\gamma^{\prime \prime}\left(s_{0}\right), \rho\left(\boldsymbol{v}_{0}\right)\right\rangle=0$, which is equivalent to $\rho\left(\boldsymbol{v}_{0}\right) \notin\left\langle\boldsymbol{t}\left(s_{0}\right), \boldsymbol{n}_{1}\left(s_{0}\right)\right\rangle_{\mathbb{R}}$. If $\boldsymbol{n}_{1}$ is a timelike vector, then $\boldsymbol{n}_{2}$ and $\boldsymbol{n}_{3}$ are spacelike vector. On the other hand, $\rho\left(\boldsymbol{v}_{0}\right) \in S_{+}^{2} \cap N \gamma$, this is a
contradiction. So $\boldsymbol{n}_{1}$ is not a timelike vector. Thus $h_{v_{0}}^{\prime}\left(s_{0}\right)=h_{v_{0}}^{\prime \prime}\left(s_{0}\right)=0$ if and only if

$$
\boldsymbol{v}_{0}= \begin{cases}\left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{2}+g_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{3}\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right) \neq 0 \\ \left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{2}\left(\lambda_{0}\right) A+g_{2}\left(\lambda_{0}\right) B\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right)=0\end{cases}
$$

or $\rho\left(v_{0}\right)=\widetilde{B}\left(s_{0}\right)$.
(3) If $k_{2} \neq 0$, then $A\left(s_{0}\right)$ is not a lightlike vector. On the other hand, $h_{v_{0}}^{\prime}\left(s_{0}\right)=$ $h_{v_{0}}^{\prime \prime}\left(s_{0}\right)=h_{v_{0}}^{\prime \prime \prime}\left(s_{0}\right)=0$ if and only if $\left\langle\left(\delta_{1} k_{1}^{2} \boldsymbol{t}+k_{1}{ }^{\prime} \boldsymbol{n}_{1}-k_{1} k_{2} \boldsymbol{n}_{2}\right)\left(s_{0}\right), \rho\left(\boldsymbol{v}_{0}\right)\right\rangle=$ $k_{1} k_{2} \delta_{2}\left(s_{0}\right)=0$, it's a contradiction. So $k_{2}\left(s_{0}\right)=0$. Thus $h_{v_{0}}^{\prime}\left(s_{0}\right)=h_{v_{0}}^{\prime \prime}\left(s_{0}\right)=$ $h_{v_{0}}^{\prime \prime \prime}\left(s_{0}\right)=0$ if and only if

$$
\boldsymbol{v}_{0}= \begin{cases}\left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{2}+g_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{3}\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right) \neq 0 \\ \left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{2}\left(\lambda_{0}\right) A+g_{2}\left(\lambda_{0}\right) B\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right)=0\end{cases}
$$

and $k_{2}\left(s_{0}\right)=0$.
(4) $h_{v_{0}}^{(4)}\left(s_{0}\right)=0$ if and only if $\left\langle\gamma^{(4)}\left(s_{0}\right), \rho\left(\boldsymbol{v}_{0}\right)\right\rangle=\left\langle\left(3 \delta_{1} k_{1} k_{1}{ }^{\prime} \boldsymbol{t}+\left(\delta_{1} k_{1}^{3}+k_{1}{ }^{\prime \prime}+\right.\right.\right.$ $\left.\left.\left.\delta_{3} k_{1} k_{2}^{2}\right) \boldsymbol{n}_{1}+\left(-2 k_{1}^{\prime} k_{2}-k_{1} k_{2}^{\prime}\right) \boldsymbol{n}_{2}+k_{1} k_{2} k_{3} \boldsymbol{n}_{3}\right)\left(s_{0}\right), \rho\left(\boldsymbol{v}_{0}\right)\right\rangle=0$, by (1), (2) and (3), $h_{v_{0}}^{\prime}\left(s_{0}\right)=h_{v_{0}}^{\prime \prime}\left(s_{0}\right)=\cdots=h_{v_{0}}^{(4)}\left(s_{0}\right)=0$ if and only if

$$
\boldsymbol{v}_{0}= \begin{cases}\left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{2}+g_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{3}\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right) \neq 0 \\ \left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{2}\left(\lambda_{0}\right) A+g_{2}\left(\lambda_{0}\right) B\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right)=0\end{cases}
$$

and $k_{2}\left(s_{0}\right)=\left\langle A^{\prime}\left(s_{0}\right), A\left(s_{0}\right)\right\rangle=0$.
Proposition 2.2. Let $\gamma: I \rightarrow \mathbb{R}_{1}^{4}$ be a unit speed spacelike curve. If $H$ is the lightcone height function of $\gamma$. Then the following conditions are equivalent:
(1) $s_{0}$ is a singularity of the lightcone Gauss map $L G_{\gamma, \eta_{0}}^{ \pm}$.
(2) $h_{v_{0}}^{\prime \prime \prime}\left(s_{0}\right)=0$ for $\boldsymbol{v}_{0}=L G_{\gamma, \eta_{0}}^{ \pm}\left(s_{0}\right)$.
(3) $k_{2}\left(s_{0}\right)=0$ and $\boldsymbol{v}_{0}=L G_{\gamma, \eta_{0}}^{ \pm}\left(s_{0}\right)$.

Proof. Let $B_{H}=\left\{\boldsymbol{v} \in S_{+}^{2} \mid h_{v}^{\prime}(s)=h_{v}^{\prime \prime}(s)=0\right\}$. By Proposition 2.1. Then $B_{H}$ can be written as $B_{H}=\left\{\boldsymbol{v} \in S_{+}^{2} \mid \boldsymbol{v}\right\}$, where

$$
\boldsymbol{v}= \begin{cases}\left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{2}+g_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{3}\right)(s), \quad k_{2}(s) \neq 0 \\ \left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{2}\left(\lambda_{0}\right) A+g_{2}\left(\lambda_{0}\right) B\right)(s), \quad k_{2}(s)=0\end{cases}
$$

or $\rho(\boldsymbol{v})=\widetilde{B}(s)$. Thus $B_{H}$ can be identified with $L G_{\gamma, \eta_{0}}^{ \pm}$from singularity theory viewpoint, which means (1) is equivalent to (2).

By Proposition 2.1, (2) is equivalent to (3).
Proposition 2.3. Let $\gamma: I \rightarrow \mathbb{R}_{1}^{4}$ be a unit speed spacelike curve. Then:
(1) The lightcone Gauss map $L G_{\gamma, \eta_{0}}^{\sigma}$ is constant if and only if $\gamma \subset \operatorname{LHP}\left(\boldsymbol{v}^{\sigma}, c^{\sigma}\right)$.
(2) Both maps, $L G_{\gamma, \eta_{0}}^{+}$and $L G_{\gamma, \eta_{0}}^{-}$, are constant if and only if $\gamma$ is a plane curve.

Here

$$
\boldsymbol{v}^{\sigma}=\left\{\begin{array}{l}
\left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{2}+g_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{3}\right)(s), \quad k_{2}(s) \neq 0 \\
\left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{2}\left(\lambda_{0}\right) A+g_{2}\left(\lambda_{0}\right) B\right)(s), \quad k_{2}(s)=0
\end{array}\right.
$$

$c^{\sigma}=\left\langle\gamma(s), \boldsymbol{v}^{\sigma}\right\rangle$ is constant, and $\sigma=+$ or - .
Proof. (1) If $L G_{\gamma, \eta_{0}}^{\sigma}$ is constant, then

$$
d\left\langle\gamma, L G_{\gamma, \eta_{0}}^{\sigma}\right\rangle=\left\langle d \gamma, L G_{\gamma, \eta_{0}}^{\sigma}\right\rangle+\left\langle\gamma, d\left(L G_{\gamma, \eta_{0}}^{\sigma}\right)\right\rangle=0
$$

Therefore $\left\langle\gamma, L G_{\gamma, \eta_{0}}^{\sigma}\right\rangle=c^{\sigma}$ is constant, which means that $\gamma \subset L H P\left(\boldsymbol{v}^{\sigma}, c^{\sigma}\right)$. Conversely, If $\gamma \subset L H P\left(\boldsymbol{v}^{\sigma}, c^{\sigma}\right)$, then $\left\langle\gamma, d L G_{\gamma, \eta_{0}}^{\sigma}\right\rangle=0$, thus $L G_{\gamma, \eta_{0}}^{\sigma}$ is constant.
(2) Since $\boldsymbol{v}^{+}$and $\boldsymbol{v}^{-}$are linearly independent, $\operatorname{LHP}\left(\boldsymbol{v}^{-}, c^{-}\right)$and $L H P\left(\boldsymbol{v}^{+}, c^{+}\right)$ intersect transversally. By (1), both maps, $L G_{\gamma, \eta_{0}}^{+}$and $L G_{\gamma, \eta_{0}}^{-}$, are constant if and only if $\gamma(s) \subset L H P\left(\boldsymbol{v}^{+}, c^{+}\right) \cap \operatorname{LHP}\left(\boldsymbol{v}^{-}, c^{-}\right)$is a plane curve.

For a unit speed spacelike curve $\gamma: I \rightarrow \mathbb{R}_{1}^{4}$, we now define extended lightcone height function $\widetilde{H}: I \times L C_{0}^{*} \cap N \gamma \rightarrow \mathbb{R}$ by $\widetilde{H}(s, \boldsymbol{v})=H(s, \widetilde{\boldsymbol{v}})-\boldsymbol{v}_{1}=\langle\gamma(s), \rho(\widetilde{\boldsymbol{v}})\rangle-$ $\boldsymbol{v}_{1}$ where $H$ is the lightcone height function of $\gamma$. For any $\boldsymbol{v}_{0}$ in $L C_{0}^{*} \cap N \gamma$, let $\widetilde{h}_{v_{0}}(s)$ denote $\widetilde{H}\left(s, \boldsymbol{v}_{0}\right)$. Then we have the following lemma.

Lemma 2.4. Let $\gamma: I \rightarrow \mathbb{R}_{1}^{4}$ be a unit speed spacelike curve with $k_{1}(s) \neq 0$. Then $\gamma$ and the hyperplane $\operatorname{LHP}\left(\rho\left(\boldsymbol{v}_{0}^{ \pm}\right), c_{0}^{ \pm}\right)$have 4-point contact at $s_{0}$ if and only if $k_{2}\left(s_{0}\right)=0$ and $\left\langle A^{\prime}\left(s_{0}\right), A\left(s_{0}\right)\right\rangle \neq 0$, where

$$
\boldsymbol{v}_{0}^{ \pm}=\left\{\begin{array}{l}
\left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{2}+g_{1}\left(\lambda_{0}\right) \boldsymbol{n}_{3}\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right) \neq 0 \\
\left(\sigma \lambda_{0} \boldsymbol{n}_{1}+f_{2}\left(\lambda_{0}\right) A+g_{2}\left(\lambda_{0}\right) B\right)\left(s_{0}\right), \quad k_{2}\left(s_{0}\right)=0
\end{array}\right.
$$

and $c_{0}^{ \pm}=\left\langle\gamma\left(s_{0}\right), \rho\left(\boldsymbol{v}_{0}^{ \pm}\right)\right\rangle$.

## 3. Unfoldings of functions of one-variable

In this section we use some general results of the singularity theory for function germs. Details can be found in [3]. Let $F:\left(\mathbb{R} \times \mathbb{R}^{r},\left(s_{0}, x_{0}\right)\right) \rightarrow \mathbb{R}$ be a function germ. We call $F$ an $r$-parameter unfolding of $f$ where $f(s)=F\left(s, x_{0}\right)$. We say that $f$ has $A_{k}$-singularity at $s_{0}$ if $f^{(p)}\left(s_{0}\right)=0$ for all $1 \leq p \leq k$ and $f^{(k+1)}\left(s_{0}\right) \neq 0$. Let $j^{(k-1)}\left(\frac{\partial F}{\partial x_{i}}\left(s, x_{0}\right)\right)\left(s_{0}\right)=\sum_{j=1}^{k-1} \alpha_{j i} s^{j}$ for $i=1, \ldots, r$, where $j^{k-1}$ denotes the $(k-1)$-jet. $F$ is called a $(p)$ versal unfolding (resp. versal) if the $(k-1) \times r$ (resp. $k \times r)$ matrix of coefficients $\left(\alpha_{j i}\right)$ (resp. $\left(\alpha_{0 i}, \alpha_{j i}\right)$ ) has rank $k-1(k-1 \leq r)$ (resp. $k(k \leq r)$ ), where $\alpha_{0 i}=\frac{\partial F}{\partial x_{i}}\left(s_{0}, x_{0}\right)$. Before proceeding further, it is convenient to introduce two important sets concerning the unfoldings. The bifurcation set of
$F$ is the set $B_{F}=\left\{x \in \mathbb{R}^{r} \left\lvert\, \frac{\partial F}{\partial s}=\frac{\partial^{2} F}{\partial^{2} s}=0\right.\right.$ at $\left.(s, x)\right\}$. The discriminant set of $F$ is the set $D_{F}=\left\{x \in \mathbb{R}^{r} \left\lvert\, F=\frac{\partial F}{\partial s}=0\right.\right.$ at $\left.(s, x)\right\}$. Then we have the following well-known result[4].

Theorem 3.1. Let $F:\left(\mathbb{R} \times \mathbb{R}^{r},\left(s_{0}, x_{0}\right)\right) \rightarrow \mathbb{R}$ be a r-parameter unfolding of $f(s)$, which has $A_{k}$-singularity at $s_{0}$.
(1) Suppose that $F$ is a $(p)$ versal unfolding.
(a) If $k=2$, then $B_{F}$ is locally diffeomorphic to $\{0\} \times \mathbb{R}^{r-1}$.
(b) If $k=3$, then $B_{F}$ is locally diffeomorphic to $C \times \mathbb{R}^{r-2}$.
(2) Suppose that $F$ is a versal unfolding.
(a) If $k=1$, then $D_{F}$ is locally diffeomorphic to $\{0\} \times \mathbb{R}^{r-1}$.
(b) If $k=2$, then $D_{F}$ is locally diffeomorphic to $C \times \mathbb{R}^{r-2}$.
(c) If $k=3$, then $D_{F}$ is locally diffeomorphic to $S W \times \mathbb{R}^{r-3}$.

Here, $S W=\left\{\left(x_{1}, x_{2}, x_{3}\right) \mid x_{1}=3 u^{4}+u^{2} v, x_{2}=4 u^{3}+2 u v, x_{3}=v\right\}$ is the swallowtail and $C=\left\{\left(x_{1}, x_{2}\right) \mid x_{1}{ }^{2}=x_{2}{ }^{3}\right\}$ is the ordinary cusp.

Theorem 3.2. Let $\gamma: I \rightarrow \mathbb{R}_{1}^{4}$ be a unit speed spacelike curve with $k_{1}(s) \neq 0, H$ the lightcone height function of $\gamma$ and $\widetilde{H}$ the extended lightcone height function.
(1) If $h(s)$ has $A_{k}$-singularity $(k=2,3)$ at $s_{0}$, then $H$ is the $(p)$ versal unfolding of $h$.

We consider the point $(s, \boldsymbol{v}) \in I \times L C_{0}^{*} \cap N_{s} \gamma$ such that $\widetilde{H}(s, \boldsymbol{v})=0$.
(2) If $\widetilde{h}(s)$ has $A_{k}$-singularity $(k=1,2,3)$ at $s_{0}$, then $\widetilde{H}$ is the versal unfolding of $\widetilde{h}$.

Proof. Let $L C_{+}^{*}=\left\{\boldsymbol{v}=\left(v_{1}, v_{2}, v_{3}, v_{4}\right) \in L C_{0}^{*} \mid v_{1}>0\right\} ; \rho(\boldsymbol{v})=\left(v_{1}, v_{2}, v_{3}, v_{4}\right) \in$ $L C_{+}^{*} ; \gamma(s)=\left(x_{1}(s), x_{2}(s), x_{3}(s), x_{4}(s)\right)$. We define a function $\bar{H}: I \times L C_{+}^{*} \rightarrow \mathbb{R}$ by $\bar{H}(s, \boldsymbol{v})=\langle\gamma(s), \rho(\boldsymbol{v})\rangle$. Since $\left.\bar{H}\right|_{I \times S_{+}^{2} \cap N \gamma}=H$, it is sufficient to verify that $\bar{H}$ is the $(p)$ versal unfolding of $\bar{h}(s)=\bar{H}\left(s, \boldsymbol{v}_{0}\right)$. In fact, since $\rho(\boldsymbol{v}) \in L C_{+}^{*}, \bar{H}(s, \boldsymbol{v})=$ $-\sqrt{v_{2}^{2}+v_{3}^{2}+v_{4}^{2}} x_{1}(s)+v_{2} x_{2}(s)+v_{3} x_{3}(s)+v_{4} x_{4}(s) \cdot \frac{\partial \bar{H}}{\partial v_{2}}=-\frac{v_{2}}{v_{1}} x_{1}(s)+x_{2}(s)$, $\frac{\partial \bar{H}}{\partial v_{3}}=x_{3}(s)-\left(\frac{v_{3}}{v_{1}}\right) x_{1}(s), \frac{\partial \bar{H}}{\partial v_{4}}=x_{4}(s)-\left(\frac{v_{4}}{v_{1}}\right) x_{1}(s)$, the 3 -jet at $s_{0}$ of $\frac{\partial \bar{H}}{\partial v_{2}}, \frac{\partial \bar{H}}{\partial v_{3}}$ and $\frac{\partial \bar{H}}{\partial v_{4}}$ are

$$
\begin{aligned}
& \left(s x_{2}^{\prime}+\frac{1}{2} s^{2} x_{2}^{\prime \prime}+\frac{1}{6} s^{3} x_{2}^{\prime \prime \prime}\right)-\left(\frac{v_{2}}{v_{1}}\right)\left(s x_{1}^{\prime}+\frac{1}{2} s^{2} x_{1}^{\prime \prime}+\frac{1}{6} s^{3} x_{1}^{\prime \prime \prime}\right), \\
& \left(s x_{3}^{\prime}+\frac{1}{2} s^{2} x_{3}^{\prime \prime}+\frac{1}{6} s^{3} x_{3}^{\prime \prime \prime}\right)-\left(\frac{v_{3}}{v_{1}}\right)\left(s x_{1}^{\prime}+\frac{1}{2} s^{2} x_{1}^{\prime \prime}+\frac{1}{6} s^{3} x_{1}^{\prime \prime \prime}\right), \\
& \left(s x_{4}^{\prime}+\frac{1}{2} s^{2} x_{4}^{\prime \prime}+\frac{1}{6} s^{3} x_{4}^{\prime \prime \prime}\right)-\left(\frac{v_{4}}{v_{1}}\right)\left(s x_{1}^{\prime}+\frac{1}{2} s^{2} x_{1}^{\prime \prime}+\frac{1}{6} s^{3} x_{1}^{\prime \prime \prime}\right) .
\end{aligned}
$$

$\bar{h}(s)$ has $A_{2}$-singularity at $s_{0}$ if and only if $\rho\left(\boldsymbol{v}_{0}\right)=\lambda\left(\boldsymbol{n}_{2}\left(s_{0}\right) \pm \boldsymbol{n}_{3}\left(s_{0}\right)\right), \lambda \neq 0$ and $k_{2}\left(s_{0}\right) \neq 0$. When $\bar{h}$ has $A_{2}$-singularity at $s_{0}$, we require matrix $\left(x_{2}^{\prime}(s)-\right.$ $\left.\left(\frac{v_{2}}{v_{1}}\right) x_{1}^{\prime}(s), x_{3}^{\prime}(s)-\left(\frac{v_{3}}{v_{1}}\right) x_{1}^{\prime}(s), x_{4}^{\prime}(s)-\left(\frac{v_{4}}{v_{1}}\right) x_{1}^{\prime}(s)\right)$ to have rank 1. In fact, Since
$\rho(\boldsymbol{v})=\lambda\left(\boldsymbol{n}_{2} \pm \boldsymbol{n}_{3}\right), k_{1} k_{2} \neq 0$, the determinant of following $3 \times 3$ matrix

$$
\begin{aligned}
& \left|\begin{array}{ccc}
x_{2}^{\prime}-\left(\frac{v_{2}}{v_{1}}\right) x_{1}^{\prime} & x_{3}^{\prime}-\left(\frac{v_{3}}{v_{1}}\right) x_{1}^{\prime} & x_{4}^{\prime}-\left(\frac{v_{4}}{v_{1}}\right) x_{1}^{\prime} \\
\frac{1}{2} x_{2}^{\prime \prime}-\left(\frac{v_{2}}{2 v_{1}}\right) x_{1}^{\prime \prime} & \frac{1}{2} x_{3}^{\prime \prime}-\left(\frac{v_{3}}{2 v_{1}}\right) x_{1}^{\prime \prime} & \frac{1}{2} x_{4}^{\prime \prime}-\left(\frac{v}{2 v_{1}}\right) x_{1}^{\prime \prime} \\
\frac{1}{6} x_{2}^{\prime \prime \prime}\left(\frac{v_{2}}{6 v_{1}}\right) x_{1}^{\prime \prime \prime} & \frac{1}{6} x_{3}^{\prime \prime \prime}-\left(\frac{v_{3}}{6 v_{1}}\right) x_{1}^{\prime \prime \prime} & \frac{1}{6} x_{4}^{\prime \prime \prime}-\left(\frac{v_{4}}{6 v_{1}}\right) x_{1}^{\prime \prime \prime}
\end{array}\right| \\
& =\frac{1}{12 v_{1}}\left|\begin{array}{cccc}
v_{1} & v_{2} & v_{3} & v_{4} \\
x_{1}^{\prime} & x_{2}^{\prime} & x_{3}^{\prime} & x_{4}^{\prime} \\
x_{1}^{\prime \prime} & x_{2}^{\prime \prime} & x_{3}^{\prime \prime} & x_{4}^{\prime \prime} \\
x_{1}^{\prime \prime \prime} & x_{2}^{\prime \prime \prime} & x_{3}^{\prime \prime \prime} & x_{4}^{\prime \prime \prime}
\end{array}\right|=\frac{k_{1}^{2} k_{2}}{12 v_{1}}\left\langle\rho(\boldsymbol{v}), \boldsymbol{n}_{3}\right\rangle
\end{aligned}
$$

is $\pm \frac{\delta_{3} \lambda k_{1}^{2} k_{2}}{12 v_{1}} \neq 0$, which means that the rank of the above matrix is 3 . Whence $\bar{H}$ is the $(p)$ versal unfolding of $\bar{h}(s)=\bar{H}\left(s, \boldsymbol{v}_{0}\right)$. Similarly for others.

By Proposition 2.1, Theorems 3.1 and 3.2, we have the following theorem.
Theorem 3.3. If $h_{v}$ and $\widetilde{h}_{w}$ have $A_{k}-\operatorname{singularity}(k=2,3)$. Then:
(1) When $k=2, B_{H}$ is locally diffeomorphic to $0 \times \mathbb{R}, D_{\widetilde{H}}$ is locally diffeomorphic to $C \times \mathbb{R}$.
(2) When $k=3, B_{H}$ is locally diffeomorphic to $C$. $D_{\widetilde{H}}$ is locally diffeomorphic to $S W$.

By Proposition 2.1 and Theorem 3.3, we get the Theorem B.

## 4. GEneric properties of spacelike curves

In this section we consider the notion of Lorentzian Monge-Taylor map of a spacelike curve in Minkowski 4 -space. Let $\gamma: I \rightarrow \mathbb{R}_{1}^{4}$ be a regular spacelike curve in Minkowski 4 -space where $I$ is an open connected subset of unit circle $S^{1}$. We now choose a smooth family of unit vectors $\boldsymbol{n}_{j}(t)$, it is pseudo perpendicular to the unit tangent vector $\boldsymbol{t}(t)$ of $\gamma$ at $t$, so $\left\|\boldsymbol{n}_{j}(t)\right\|=1$ and $\left\langle\boldsymbol{n}_{j}(t), \boldsymbol{t}(t)\right\rangle=0$ for all $t \in I$. Such $\boldsymbol{n}_{j}(t)$ can be obtained as following: consider the smooth map $\boldsymbol{t}: I \rightarrow S_{1}^{3}$ which takes $t$ to the unit tangent vector $\boldsymbol{t}(t)$, if $V$ is a vector in $S_{1}^{3}$, we can obtain the vector field $\boldsymbol{n}_{j}(t)$ by pseudo-orthogonally projecting $V$ onto each of the pseudo normal space and normalizing. Thus $\boldsymbol{n}_{j}(t)=\frac{V-\langle V, \boldsymbol{t}(t)\rangle \boldsymbol{t}(t)}{\|V-\langle V, \boldsymbol{t}(t)\rangle \boldsymbol{t}(t)\|} . \boldsymbol{n}_{i}(t)$ is obtained similarly:

$$
\boldsymbol{n}_{i}(t)=\frac{W-\left\langle W, \boldsymbol{n}_{j}(t)\right\rangle \boldsymbol{n}_{j}(t)-\langle W, \boldsymbol{t}(t)\rangle \boldsymbol{t}(t)}{\left\|W-\left\langle W, \boldsymbol{n}_{j}(t)\right\rangle \boldsymbol{n}_{j}(t)-\langle W, \boldsymbol{t}(t)\rangle \boldsymbol{t}(t)\right\|},
$$

where $W \in H^{3}\left(H^{3}=\left\{p \in \mathbb{R}_{1}^{4} \mid\langle p, p\rangle=-1\right\}\right)$. Let $\boldsymbol{n}_{k}(t)=\boldsymbol{t}(t) \wedge \boldsymbol{n}_{j}(t) \wedge \boldsymbol{n}_{i}(t)$. We use the pseudo perpendicular lines spanned by $\boldsymbol{t}(t), \boldsymbol{n}_{j}(t), \boldsymbol{n}_{i}(t), \boldsymbol{n}_{k}(t)$ as axes at $\gamma(t)$ with the unit points on the axes corresponding to the four given vectors.

Note that the curve $\gamma$ is not necessarily unit speed with $\gamma\left(t_{0}\right)=0$. We can write $\gamma(I)$ locally as $\left(g_{t}(\xi), \xi, f_{t}(\xi), l_{t}(\xi)\right.$ with $j^{1} f_{t}(0)=j^{1} g_{t}(0)=j^{1} l_{t}(0)=0$. If $V_{k}$ denotes the space of polynomials in $\xi$ of degree $\geq 2$ and $\leq k$, then we have a map, the Monge-Taylor map for the spacelike curve $\gamma, \mu_{\gamma}: I \rightarrow V_{k} \times$ $V_{k} \times V_{k}$ by $\mu_{\gamma}(t)=\left(j^{k} f_{t}(0), j^{k} g_{t}(0), j^{k} l_{t}(0)\right) . V_{k} \times V_{k} \times V_{k}$ can be identified with $\mathbb{R}_{1}^{k-1} \times \mathbb{R}_{1}^{k-1} \times \mathbb{R}_{1}^{k-1}=\mathbb{R}_{1}^{3 k-3}$ via the coordinate $\left(a_{2}, \ldots, a_{k}, b_{2}, \ldots, b_{k}, c_{2}, \ldots, c_{k}\right)$. Of course $\mu_{\gamma}$ depends rather heavily on our choice of unit normals $\boldsymbol{n}_{j}(t)$ and $\boldsymbol{n}_{i}(t)$. Here $a_{i}=\frac{f_{t}^{(i)}(0)}{i!}, b_{i}=\frac{g_{t}^{(i)}(0)}{i!}, c_{i}=\frac{l_{t}^{(i)}(0)}{i!}(2 \leq i \leq k)$, that is $V_{k} \times V_{k} \times$ $V_{k}=\left\{\left(a_{2} \xi^{2}+a_{3} \xi^{3}+\cdots+a_{k} \xi^{k}\right),\left(b_{2} \xi^{2}+b_{3} \xi^{3}+\cdots+b_{k} \xi^{k}\right),\left(c_{2} \xi^{2}+c_{3} \xi^{3}+\cdots+\right.\right.$ $\left.\left.c_{k} \xi^{k}\right)\right\}$. Let $P_{k}$ denote the set of maps $\psi: \mathbb{R}_{1}^{4} \rightarrow \mathbb{R}_{1}^{4}$ of the form $\psi(x, y, z, w)=$ $\left(\left(\psi_{1}(x, y, z, w), \psi_{2}(x, y, z, w), \psi_{3}(x, y, z, w), \psi_{4}(x, y, z, w)\right)\right.$, where $\psi_{i}(x, y, z, w)$ is a polynomial in $x, y, z, w$ of degree $\leq k$. An element $\psi \in P_{k}$ is determined by the coefficients of the various monomials $x^{i} y^{j} z^{m} w^{n}$ occurring in $\psi_{1}, \psi_{2}, \psi_{3}, \psi_{4}$. There are altogether $\frac{(k+4)!}{4!k!}$ monomials $x^{i} y^{j} z^{m} w^{n}$ of degree $\leq k$, so that $P_{k}$ can be thought as a Minkowski space $\mathbb{R}_{1}^{\frac{(k+4)!}{1!k!}}$. It is this space that will provide the required deformations of the curve.

To simplify matters we now assume that the curve $\gamma$ is compact, that is $I=S^{1}$. The identity map $1_{\mathbb{R}_{1}^{4}}: \mathbb{R}_{1}^{4} \rightarrow \mathbb{R}_{1}^{4}$ is of course an element of $P_{k}(k \geq 1)$. By using compactness of $\gamma$, it is easy to see that there is an open neighborhood $U$ of $1_{\mathbb{R}_{1}^{4}}$ in $P_{k}$ with the property that if $\psi \in U$, then the linear map $T \psi(\gamma(t)): \mathbb{R}_{1}^{4} \rightarrow \mathbb{R}_{1}^{4} ; \boldsymbol{v} \longmapsto D \psi(\gamma(t)) \cdot \boldsymbol{v}$ satisfies that it takes a timelike vector(resp. a spacelike vector) to a timelike vector (resp. a spacelike vector), where $D \psi(\gamma(t))$ denotes the derivative of $\psi$ at $\gamma(t)$. If we deform the original curve by the map $\psi$, then we can also obtain the required two new smooth family of normal vectors $\boldsymbol{n}_{i \psi(t)}, \boldsymbol{n}_{j \psi(t)}$ as follows. Since the map $\psi: \mathbb{R}_{1}^{4} \rightarrow \mathbb{R}_{1}^{4}$ is a diffeomorphism on some open set containing $\gamma(I)$, vectors $\boldsymbol{n}_{j}(t), \boldsymbol{n}_{i}(t)$ will be sent to some new vectors $D \psi(\gamma(t)) \boldsymbol{n}_{j}(t), D \psi(\gamma(t)) \boldsymbol{n}_{i}(t)$, which will be neither zero nor tangent to $\psi \circ \gamma$ at $t$. Pseudo-orthogonally project $D \psi(\gamma(t)) \boldsymbol{n}_{j}(t), D \psi(\gamma(t)) \boldsymbol{n}_{i}(t)$ onto the pseudonormal space to $\psi \circ \gamma$ at $t$ and normalize, then we get $\boldsymbol{n}_{j \psi}(t)=\frac{D \psi(\gamma(t)) \boldsymbol{n}_{j}(t)-\left\langle D \psi(\gamma(t)) \boldsymbol{n}_{j}(t), \boldsymbol{t}_{\psi}\right\rangle \boldsymbol{t}_{\psi}}{\left.\| D \psi(\gamma(t)) \boldsymbol{n}_{j}(t)-\left\langle D \psi(\psi(t)) \boldsymbol{n}_{j} t\right), \boldsymbol{t}_{\psi}\right\rangle \boldsymbol{t}_{\psi} \|},\left\langle\boldsymbol{n}_{j \psi}(t), \boldsymbol{n}_{\boldsymbol{j} \psi}(t)\right\rangle=1$,

$$
\boldsymbol{n}_{i \psi}(t)=\frac{D \psi(\gamma(t)) \boldsymbol{n}_{i}(t)-\left\langle D \psi(\gamma(t)) \boldsymbol{n}_{i}(t), \boldsymbol{n}_{j \psi}\right\rangle \boldsymbol{n}_{j \psi}-\left\langle D \psi(\gamma(t)) \boldsymbol{n}_{i}(t), \boldsymbol{t} \psi\right\rangle \boldsymbol{t} \psi}{\| D \psi(\gamma(t)) \boldsymbol{n}_{i}(t)-\left\langle D \psi(\gamma(t)) \boldsymbol{n}_{i}(t), \boldsymbol{n}_{j \psi}\right\rangle \boldsymbol{n}_{j \psi}-\left\langle D \psi(\gamma(t)) \boldsymbol{n}_{i}(t), \boldsymbol{t} \psi \|\right.}
$$

and $\left\langle\boldsymbol{n}_{i \psi}(t), \boldsymbol{n}_{i \psi}(t)\right\rangle=-1$, where $\boldsymbol{t}_{\psi}$ denotes the tangent vector of the curve $\psi \circ \gamma$ at $t$. Assuming as before that $I=S^{1}$, we choose an open neighborhood $U$ of $1_{I} \in P_{k}$ consisting of polynomial maps which map an open set containing $\gamma\left(S^{1}\right)$ diffeomorphic to its image. We have now shown that there is a smooth map

$$
\mu: S^{1} \times U \rightarrow V_{k} \times V_{k} \times V_{k}
$$

defined by $\mu(-, \psi)=$ Monge-Taylor map for the curve $\psi \circ \gamma$ using the family of pseudo-normal vectors $\boldsymbol{n}_{j \psi}(t)$ and $\boldsymbol{n}_{i \psi}(t)$. By the same arguments as in the proof of Theorem 9.9 in [4], we have the following theorem.

Theorem 4.1. Let $Q$ be a manifold in $V_{k} \times V_{k} \times V_{k}=\mathbb{R}_{1}^{3 k-3}$. For some open set $U_{1} \subset U$ containing identity map, the map $\mu: S^{1} \times U_{1} \rightarrow V_{k} \times V_{k} \times V_{k}$ defined by $\mu(t, \psi)=\mu_{\psi \circ \gamma}(t)$ is transverse to $Q$ (In fact, $\mu$ is a submersion, so that $Q$ does not enter the argument at all).

A straightforward computation shows the following lemma. The computations are rather long and tedious, so we omit the details.

Lemma 4.2. Let $\gamma$ be a unit speed spacelike curve defined by
$\gamma(t)=\left(g_{t}(\xi), \xi, f_{t}(\xi), l_{t}(\xi)\right)=\left(b_{2} \xi^{2}+b_{3} \xi^{3}+\cdots, \xi, a_{2} \xi^{2}+a_{3} \xi^{3}+\cdots, c_{2} \xi^{2}+c_{3} \xi^{3}+\cdots\right)$
with $\xi\left(t_{0}\right)=0$ and $k_{1}\left(t_{0}\right) \neq 0$. We let $N_{i j}=i!j!\left(a_{i} a_{j}-b_{i} b_{j}+c_{i} c_{j}\right)$. Then:
(1) $f_{1}\left(a_{2}, a_{3}, b_{2}, b_{3}, c_{2}, c_{3}\right)=0$ at $t_{0}$ if and only if $k_{2}\left(t_{0}\right)=0$, where $f_{1}=$ $N_{32}^{2}+N_{22}^{3}-N_{33} N_{22}$.
(2) $f_{2}\left(a_{2}, a_{3}, a_{4}, b_{2}, b_{3}, b_{4}, c_{2}, c_{3}, c_{4}\right)=0$ at $t_{0}$ if and only if $\left\langle A^{\prime}, A\right\rangle\left(t_{0}\right)=0$, where $f_{2}=N_{22}^{3} N_{34}-2 N_{32} N_{22}^{2} N_{33}+N_{22} N_{32}\left[3 N_{32}^{2}-N_{22}\left(N_{42}+N_{33}\right)-N_{22}^{3}\right]-$ $N_{32} N_{22}^{3} N_{31}-N_{22}^{2} N_{32} N_{42}+2 N_{32}^{3} N_{22}-N_{32} N_{22}\left[3 N_{32}^{2}-N_{22}\left(N_{42}+N_{33}\right)-N_{22}^{3}\right]+$ $N_{32}^{2} N_{22}^{2} N_{21}-N_{22}^{4} N_{41}+2 N_{32} N_{22}^{3} N_{31}-N_{22}^{2} N_{21}\left[3 N_{32}^{2}-N_{22}\left(N_{42}+N_{33}\right)-N_{22}^{3}\right]+$ $N_{32} N_{22}^{4} N_{11}$.

Here, $\xi$ is the coordinate along the $\boldsymbol{t}$-direction, $f_{t}(\xi)$ is the coordinate along the $\boldsymbol{n}_{j}$-direction, $g_{t}(\xi)$ is the coordinate along the $\boldsymbol{n}_{i}$-direction, $l_{t}(\xi)$ is the coordinate along the $\boldsymbol{n}_{k}$-direction.

Lemma 4.3. We consider smooth maps $\rho_{i}: V_{4} \times V_{4} \times V_{4}=\mathbb{R}^{9} \rightarrow \mathbb{R}(i=1,2)$ given by

$$
\left\{\begin{aligned}
\rho_{1}= & N_{32}^{2}+N_{22}^{3}-N_{33} N_{22} ; \\
\rho_{2}= & N_{22}^{3} N_{34}-2 N_{32} N_{22}^{2} N_{33}+N_{22} N_{32}\left[3 N_{32}^{2}-N_{22}\left(N_{42}+N_{33}\right)-N_{22}^{3}\right] \\
& -N_{32} N_{22}^{3} N_{31}-N_{22}^{2} N_{32} N_{42}+2 N_{32}^{3} N_{22} \\
& -N_{32} N_{22}\left[3 N_{32}^{2}-N_{22}\left(N_{42}+N_{33}\right)-N_{22}^{3}\right]+N_{32}^{2} N_{22}^{2} N_{21} \\
& -N_{22}^{4} N_{41}+2 N_{32} N_{22}^{3} N_{31}-N_{22}^{2} N_{21}\left[3 N_{32}^{2}-N_{22}\left(N_{42}+N_{33}\right)-N_{22}^{3}\right] \\
& +N_{32} N_{22}^{4} N_{11} .
\end{aligned}\right.
$$

Then the set $Q_{1}=\left\{\left(a_{2}, a_{3}, a_{4}, b_{2}, b_{3}, b_{4}, c_{2}, c_{3}, c_{4}\right) \in \mathbb{R}^{9} \mid \rho_{1}=\rho_{2}=0\right\}$ is a codimension two submanifold in $\mathbb{R}^{9}$.

An ordinary(resp. degenerate) vertex is a point $p=\gamma\left(t_{0}\right)$ of a spacelike curve $\gamma$, for which there exists a hyperplane having 4-(resp. at least 5 -) point contact with the curve for $t=t_{0}$. We say $\gamma$ has a vertex at $t_{0}$, or at $p$.

We now give the proof of Theorem A.

Proof. Suppose we are given any regular spacelike curve $\gamma: S^{1} \rightarrow \mathbb{R}_{1}^{4}$. Applying Theorem 4.1 to the map $\mu$ with $Q$ the submanifold of degenerate vertexes proves that a dense set of curves have only ordinary vertexes. Further taking $k=4$ and $Q$ to be the submanifold of $f_{1}(t)=0$. Then the vertexes of $\psi \circ \gamma$ correspond to points $t$ with $\mu(t, \psi) \in Q$. However, again by Theorem 4.1, we know that for a dense set of $\psi \in U_{1}$ the map $\mu(-, \psi)=\nu$ is transverse to $Q$. Consequently the set $\nu^{-1}(Q)$ of the ordinary vertexes is finite for a dense set of $\psi \in U_{1}$, whence the result.

We now have to prove that these properties are open. We first show that this is so for the property of only having ordinary vertexes. Let $Q$ denote the set $f_{1}=f_{2}=0$ in $V_{4} \times V_{4} \times V_{4}$ (degenerate vertexes). Thus $Q$ is closed. Let $\widetilde{\gamma}: S^{1} \times U \rightarrow \mathbb{R}_{1}^{4}$ be a family of curves with $\widetilde{\gamma}_{0}=\widetilde{\gamma}(-, 0)$ having only ordinary vertexes. Let $\mu: S^{1} \times U \rightarrow V_{k} \times V_{k} \times V_{k}$ be the corresponding family of MongeTaylor map. Then the compactness of $S^{1}$, together with the fact that $\mu_{0}=$ $\mu(-, 0)$ is transverse to $Q$ (misses $Q$ in fact), implies by [4, Proposition 8.23] that $\mu\left(S^{1} \times\{u\}\right)$ misses $Q$ for $u$ in some open neighborhood $U^{\prime}$ of 0 . Hence nearby curves $\widetilde{\gamma}_{u}$ in the family also possess no degenerate vertex.

It remains to show that, in fact, the property of having finitely many ordinary vertexes and no degenerate vertex is open. First note that, if $\gamma$ has an ordinary vertex at $t \in S^{1}$, then the image of the map $\mu: S^{1} \rightarrow V_{4} \times V_{4} \times V_{4}$ meets the submanifold of $f_{1}=0$ at $\mu(t)$ and is transverse to this submanifold.

Let $\widetilde{\gamma}: S^{1} \times U \rightarrow \mathbb{R}_{1}^{4}$ be a family of curves with $\widetilde{\gamma}_{0}$ having finitely many ordinary vertexes, and no degenerate vertex, so that $\mu(-, 0)$ is transverse to the submanifold of $f_{1}=0$. Since transversality is an open condition when the source is compact and the relevant submanifold are closed[4, Proposition 8.23], it follows that $\nu=\mu(-, u): S^{1} \rightarrow V_{4} \times V_{4} \times V_{4}$ will also transverse to that submanifold for all $u$ in some neighborhood $U_{1}$ of $0 \in U$. Consequently, if $Q$ is that submanifold, the set $\nu^{-1}(Q)$ of ordinary vertexes of $\widetilde{\gamma}_{u}$ is finite and there is no degenerate vertex. This proves the result.

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